ADITYA L1 MISSION: SCIENCE & TECHNOLOGY

NEWS: Aditya-L1 Mission Captures First-Ever Image of Solar Flare 'Kernel'

WHAT'S IN THE NEWS?

The Indian Space Research Organisation (ISRO) announced that its Aditya-L1 mission has successfully captured the first-ever image of a solar flare 'kernel' using the Solar Ultraviolet Imaging Telescope (SUIT) payload. This marks a significant advancement in understanding solar flare dynamics and energy transfer in the Sun's atmosphere.

Key Observations from Aditya-L1's SUIT Payload

- The **SUIT payload** aboard **Aditya-L1** captured the image in the **lower solar atmosphere**, specifically in the **photosphere and chromosphere**.
- The solar flare was observed in the Near Ultraviolet (NUV) wavelength range of 200-400 nm, an unprecedented achievement in solar research.
- The flare detected was an X6.3-class solar flare, one of the most intense solar eruptions recorded.
- The observation confirmed that **energy released from the flare spreads through different layers of the Sun's atmosphere**, validating existing solar flare models.
- The localized brightening in the lower solar atmosphere was found to correspond with an increase in plasma temperature in the solar corona.
- These findings reinforce theories linking flare energy deposition to temperature evolution in the Sun's outer layers.

Overview of Aditya-L1 Mission

- Aditya-L1 is India's first space-based observatory dedicated to studying the Sun and solar activity from space.
- It is designed to monitor **solar emissions**, **flares**, **and space weather phenomena** that impact Earth's climate and satellite communications.

Key Payloads of Aditya-L1

Aditya-L1 is equipped with seven advanced payloads to study various solar phenomena:

- 1. Visible Emission Line Coronagraph (VELC) Observes the corona and studies coronal mass ejections (CMEs).
- 2. Solar Ultraviolet Imaging Telescope (SUIT) Captures high-resolution images of the photosphere and chromosphere in the Near Ultraviolet (NUV) spectrum.
- 3. Solar Low Energy X-ray Spectrometer (SoLEXS) Monitors soft X-ray emissions from solar flares.
- 4. Aditya Solar wind Particle Experiment (ASPEX) Studies solar wind properties and its interactions with Earth's magnetosphere.
- 5. **High Energy L1 Orbiting X-ray Spectrometer (HEL1OS)** Detects high-energy X-rays from solar flares.

- 6. **Plasma Analyser Package for Aditya (PAPA)** Analyzes the solar wind's composition and energy distribution.
- 7. Advanced Tri-axial High-Resolution Digital Magnetometers Measures magnetic field variations in space.

Understanding Lagrange Points and Aditya-L1's Orbit

- The "L1" in Aditya-L1 refers to the Lagrange Point 1, a special position in space where the gravitational forces of the Sun and Earth balance the centrifugal force of a satellite.
- Lagrange points are important for space missions as they **reduce fuel consumption** required to maintain a stable orbit.

Types of Lagrange Points in the Sun-Earth System

- 1. L1 (Lagrange Point 1):
 - Positioned between the **Sun and Earth**.
 - Provides an **uninterrupted view of the Sun**, making it ideal for solar observation missions like **Aditya-L1** and **SOHO** (Solar and Heliospheric **Observatory**).
- 2. L2 (Lagrange Point 2):
 - Located on the **opposite side of the Earth from the Sun**.
 - Used by deep-space observatories such as **James Webb Space Telescope** (**JWST**).
- 3. L3 (Lagrange Point 3):
 - Lies directly behind the Sun, opposite to Earth's orbit.
 - Not widely used due to communication difficulties.
- 4. L4 and L5 (Stable Lagrange Points):
 - Form equilateral triangles with the Sun and Earth.
 - Natural accumulation points for small celestial bodies (e.g., asteroids known as **Trojans**).

Solar Flares and Their Impact

What is a Solar Flare?

- A solar flare is a sudden, intense burst of energy originating from the Sun's atmosphere, caused by the reconnection of magnetic fields.
- It releases energy in the form of light, radiation, and high-energy particles, which can disrupt satellite communications, power grids, and navigation systems on Earth.

What is a Solar Flare 'Kernel'?

- In solar physics, the 'kernel' of a solar flare refers to:
 - The **brightest and most concentrated region** observed in the **lower layers of the Sun's atmosphere** during a flare.

- It is a site of rapid heating and plasma upflow during the rise phase of a solar flare.
- Studying kernels helps scientists understand the energy transfer process from the Sun's core to its outer layers.



Classification of Solar Flares

Solar flares are categorized based on their **peak X-ray flux** in the **1 to 8 Angstrom range**, as measured by the **Geostationary Operational Environmental Satellites (GOES)**.

Types of Solar Flares

- 1. A-Class Flares (Weakest):
 - Peak flux: **10⁻⁸ W/m²**.
 - Minimal impact, barely distinguishable from background solar radiation.
- 2. B-Class Flares:
 - Peak flux: **10**⁻⁷ **W**/**m**².
 - Ten times stronger than A-class but still has negligible impact on Earth.
- 3. C-Class Flares:
 - Peak flux: 10⁻⁶ W/m².
 - Considered minor flares with little effect on Earth.
- 4. M-Class Flares (Moderate intensity):
 - Peak flux: 10⁻⁵ W/m².

- Can cause temporary radio blackouts in Earth's polar regions.
- Sometimes associated with small radiation storms.
- 5. X-Class Flares (Most powerful):
 - Peak flux: 10⁻⁴ W/m² or higher.
 - Can trigger planet-wide radio blackouts and long-lasting radiation storms.
 - Each numerical increment represents a tenfold increase in energy output (e.g., X2 flare is twice as powerful as an X1 flare).

Significance of X6.3-Class Flare Observed by Aditya-L1

- The captured X6.3-class solar flare is one of the strongest flares recorded.
- Such flares have the potential to cause:
 - Severe space weather disruptions affecting satellite operations.
 - **Increased radiation hazards for astronauts** and aircraft flying over polar regions.
 - Geomagnetic storms that may impact Earth's power grids and communication networks.

Significance of Aditya-L1's Discoveries

- First-ever direct imaging of a solar flare kernel in the NUV spectrum.
- Enhances understanding of **solar energy transfer mechanisms** from the Sun's **core to its outer atmosphere**.
- Helps improve space weather forecasting, reducing risks to satellites, astronauts, and Earth's communication systems.
- Strengthens India's position in **solar research and space exploration**, contributing to global scientific advancements.

Source: <u>https://www.thehindu.com/sci-tech/science/aditya-l1-payload-captures-the-first-ever-image-of-a-solar-flare-kernel/article69274469.ece</u>